

Assignments to Condensed Matter Theory I

Sheet 11

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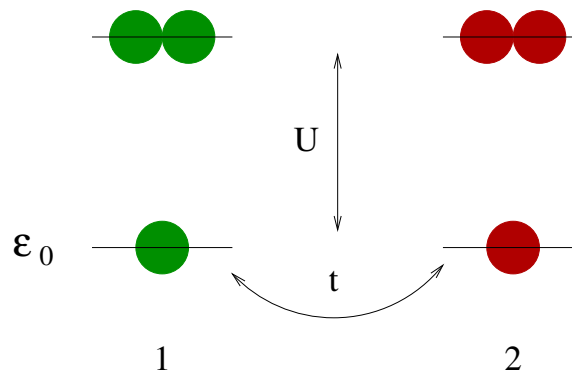
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Problem set: Electron-Electron Interaction (II)

11.1. Double site Hubbard model

The Hubbard Hamiltonian for a two site system reads explicitly:

$$H = \epsilon_0 \left(c_{1\uparrow}^\dagger c_{1\uparrow} + c_{1\downarrow}^\dagger c_{1\downarrow} + c_{2\uparrow}^\dagger c_{2\uparrow} + c_{2\downarrow}^\dagger c_{2\downarrow} \right) + t \left(c_{1\uparrow}^\dagger c_{2\uparrow} + c_{2\downarrow}^\dagger c_{1\downarrow} + c_{2\uparrow}^\dagger c_{1\uparrow} + c_{1\downarrow}^\dagger c_{2\downarrow} \right) + U \left(c_{1\uparrow}^\dagger c_{1\uparrow} c_{1\downarrow}^\dagger c_{1\downarrow} + c_{2\uparrow}^\dagger c_{2\uparrow} c_{2\downarrow}^\dagger c_{2\downarrow} \right)$$



- (a) Calculate the two particle eigenenergies analytically. Treat the case of parallel and antiparallel spin separately. Plot the results as a function of U/t .

Hint: For the antiparallel case consider the basis of the corresponding Hilbert space:

$$c_{1\uparrow}^\dagger c_{1\downarrow}^\dagger |0\rangle, c_{2\uparrow}^\dagger c_{2\downarrow}^\dagger |0\rangle, c_{1\uparrow}^\dagger c_{2\downarrow}^\dagger |0\rangle, c_{2\uparrow}^\dagger c_{1\downarrow}^\dagger |0\rangle$$

Calculate the matrix elements of H in this basis and diagonalize the resulting 4×4 matrix.

- (b) Calculate the ground state in the Hartree-Fock approximation and compare it with the exact result of point (a).

11.2. Homogeneous electron gas in the Hartree-Fock approximation

Show that the following identities hold by transforming the sums over \mathbf{k}' and k into integrals and evaluate them.

(a)

$$\frac{1}{V} \sum_{\substack{\mathbf{k}' \\ (k' \leq k_F)}} \frac{1}{|\mathbf{k} - \mathbf{k}'|^2} = \frac{1}{2\pi^2} k_F \left[\frac{1}{2} + \frac{1 - (k/k_F)^2}{4(k/k_F)} \ln \left| \frac{1 + k/k_F}{1 - k/k_F} \right| \right]$$

(b)

$$\sum_{k < k_F} \frac{k_F^2 - k^2}{kk_F} \ln \left| \frac{k + k_F}{k - k_F} \right| = \frac{V}{6\pi^2} k_F^3.$$

11.3. [Kür] Thomas Fermi limit in atomic physics

Let us consider an atom with a positively charged nucleus with a charge $Z|e|$ (where e is the electronic charge taken with the sign) and its surrounding electron shells. In the Hartree limit the potential felt by each electron at position \mathbf{r} is given by the expression:

$$V(\mathbf{r}) = \int d\mathbf{r}' \frac{e^2}{|\mathbf{r} - \mathbf{r}'|} n(\mathbf{r}') - \frac{Ze^2}{r} \quad (1)$$

where $n(\mathbf{r})$ is the density of electrons. In the Thomas-Fermi approximation one can assume that the potential is varying so slowly that can be considered constant around a given point \mathbf{r} . Then it is allowed to introduce a local homogeneous electron gas for the region around the point \mathbf{r} .

(a) Justify that, from the previous assumption, it follows that

$$n(\mathbf{r}) = \frac{[2m(\epsilon_F - V(\mathbf{r}))]^{3/2}}{3\pi^2}$$

where ϵ_F is the energy of the highest occupied energy level.

(b) Justify that for a neutral atom $\epsilon_F = 0$.

(c) Prove that Eq. (1) is equivalent to the Poisson equation and derive from that and using the result of point (b) the Thomas-Fermi equation:

$$-\frac{3\pi}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial V(r)}{\partial r} \right) = 4e^2 [2m(\epsilon_F - V(\mathbf{r}))]^{3/2} \quad (2)$$

Hint: The charge distribution and the associated potential can be considered spherically symmetric.

(d) Show that, by introducing the Bohr radius a_0 as length unit,

$$V(r) = -\frac{Ze^2}{r} \Phi(x), \quad r = Z^{-1/3} b x, \quad b = \frac{1}{2} \left(\frac{3\pi}{4} \right)^{2/3} a_0, \quad a_0 = \frac{1}{me^2}$$

Eq. (2) becomes the following differential equation for the dimensionless potential $\Phi(x)$:

$$\frac{d^2}{dx^2} \Phi(x) = x^{-1/2} \Phi^{3/2}(x) \quad (3)$$

with the boundary conditions

$$\Phi(0) = 1, \lim_{x \rightarrow \infty} \Phi(x) = 0 \quad (4)$$

Note: We have assumed $\hbar = 1$ and $4\pi\epsilon_0 = 1$.

(e) Solve Eq. (3) numerically and plot the functions $\Phi(x)$, $V(r)$ and $4\pi r^2 n(r)$.

SOLUTIONS OF THE EXERCISES FOR THE SHEET 11 OF QTKM1

11.1 Double site Hubbard model

$$\begin{aligned}
 H = & \varepsilon_0 (c_{1\uparrow}^\dagger c_{2\uparrow} + c_{1\downarrow}^\dagger c_{2\downarrow} + c_{2\uparrow}^\dagger c_{1\uparrow} + c_{2\downarrow}^\dagger c_{1\downarrow}) + \\
 & t (c_{1\uparrow}^\dagger c_{2\uparrow} + c_{2\downarrow}^\dagger c_{1\downarrow} + c_{2\uparrow}^\dagger c_{1\uparrow} + c_{1\downarrow}^\dagger c_{2\downarrow}) \\
 & + U (c_{1\uparrow}^\dagger c_{1\uparrow} c_{1\downarrow}^\dagger c_{1\downarrow} + c_{2\uparrow}^\dagger c_{2\uparrow} c_{2\downarrow}^\dagger c_{2\downarrow})
 \end{aligned}$$

a) Two particles with parallel spin can be inserted in the system only in the form:

$$c_{1\uparrow}^\dagger c_{2\uparrow}^\dagger |0\rangle \quad \text{or} \quad c_{1\downarrow}^\dagger c_{2\downarrow}^\dagger |0\rangle$$

Note that we assume the order of the particles to be

$$1\uparrow \ 2\uparrow \ 1\downarrow \ 2\downarrow$$

PARALLEL

$$c_{1\uparrow}^\dagger c_{2\uparrow}^\dagger |0\rangle \equiv |+, +\rangle$$

$$c_{1\downarrow}^\dagger c_{2\downarrow}^\dagger |0\rangle \equiv |-, -\rangle$$

$$H |+, +\rangle = 2\varepsilon_0 |+, +\rangle$$

$$H |-, -\rangle = 2\varepsilon_0 |-, -\rangle$$

ANTIPARALLEL

The space of the antiparallel case is spanned by the 4 vectors:

$$|z, 0\rangle \equiv c_{1\uparrow}^\dagger c_{1\downarrow}^\dagger |0\rangle, \quad |0, z\rangle \equiv c_{2\uparrow}^\dagger c_{2\downarrow}^\dagger |0\rangle;$$

$$|+, -\rangle \equiv c_{1\uparrow}^\dagger c_{2\downarrow}^\dagger |0\rangle, \quad |-, +\rangle \equiv c_{2\uparrow}^\dagger c_{1\downarrow}^\dagger |0\rangle$$

$$H|2,0\rangle = 2\varepsilon_0|2,0\rangle + t(|+,-\rangle + |- ,+\rangle) + U|2,0\rangle$$

$$H|0,2\rangle = 2\varepsilon_0|0,2\rangle + t(|+,-\rangle + |- ,+\rangle) + U|0,2\rangle$$

$$H|+,-\rangle = 2\varepsilon_0|+,-\rangle + t(|0,2\rangle + |2,0\rangle)$$

$$H|- ,+\rangle = 2\varepsilon_0|- ,+\rangle + t(|0,2\rangle + |2,0\rangle)$$

We choose the basis with the order :

$$|2,0\rangle, |0,2\rangle, |+,-\rangle, |- ,+\rangle$$

The matrix representation of H reads:

$$H = \begin{bmatrix} 2\varepsilon_0 + U & 0 & t & t \\ 0 & 2\varepsilon_0 + U & t & t \\ t & t & 2\varepsilon_0 & 0 \\ t & t & 0 & 2\varepsilon_0 \end{bmatrix}$$

$\det |H - \lambda \mathbb{1}| = 0$ in the eq. in λ to find the eigenvalues. We

define $\mu = \lambda - 2\varepsilon_0$

$$0 = \det \begin{vmatrix} U-\mu & 0 & t & t \\ 0 & U-\mu & t & t \\ t & t & -\mu & 0 \\ t & t & 0 & -\mu \end{vmatrix} = (U-\mu) \det \begin{vmatrix} U-\mu & t & t \\ t & -\mu & 0 \\ t & 0 & -\mu \end{vmatrix} + t \det \begin{vmatrix} 0 & U-\mu & t \\ t & t & 0 \\ t & t & -\mu \end{vmatrix}$$

$$- t \det \begin{vmatrix} 0 & U-\mu & t \\ t & t & -\mu \\ t & t & 0 \end{vmatrix} =$$

$$= (U-\mu) [(U-\mu)\mu^2 + 2\mu t^2] + 2t [t^3 - t^3 + \mu t(U-\mu)] =$$

$$= (U-\mu) [(U-\mu)\mu^2 + 4\mu t^2] = (U-\mu)\mu [(U-\mu)\mu + 4t^2] = \mu(U-\mu) [-\mu^2 + U\mu + 4t^2]$$

$$\mu_1 = 0 \quad \lambda_1 = 2\varepsilon_0$$

$$\mu_2 = U \quad \lambda_2 = 2\varepsilon_0 + U$$

$$\mu_{3,4} = \frac{+U \pm \sqrt{U^2 + 16t^2}}{2} \quad \lambda_{3,4} = 2\varepsilon_0 + \frac{U}{2} \left(1 \pm \sqrt{1 + \frac{16t^2}{U^2}} \right)$$

The ground state energy is given in any case by

$$\lambda_3 = 2\varepsilon_0 + \frac{U}{2} \left(1 - \sqrt{1 + \frac{16t^2}{U^2}} \right)$$

If $t \ll U$

$$\lambda_{3,4} \approx 2\varepsilon_0 - \frac{4t^2}{U}, \quad 2\varepsilon_0 + U + \frac{t^2}{U}$$

If $t \gg U$

$$\lambda_{3,4} \approx 2\varepsilon_0 + \frac{U}{2} \left(1 \pm \frac{4|t|}{U} \sqrt{1 + \frac{U^2}{16t^2}} \right)$$

$$= 2\varepsilon_0 + 2|t| \left(\frac{U}{4|t|} \pm \sqrt{1 + \frac{U^2}{16t^2}} \right) \sim 2\varepsilon_0 \pm 2|t| + \frac{U}{2}$$

b) The ground state in Hartree-Fock approximation can be calculated starting with:

$$H_{\text{HF}} = \varepsilon_0 \sum_{i\sigma} c_{i\sigma}^\dagger c_{i\sigma} + t \sum_{\sigma} \left(c_{1\sigma}^\dagger c_{2\sigma} + c_{2\sigma}^\dagger c_{1\sigma} \right)$$

$$+ U \left(\bar{n}_{1\uparrow} c_{1\uparrow}^\dagger c_{1\uparrow} + c_{1\uparrow}^\dagger c_{1\uparrow} \bar{n}_{1\uparrow} - \bar{n}_{1\uparrow} \bar{n}_{1\uparrow} + \right.$$

$$\left. + \bar{n}_{2\uparrow} c_{2\uparrow}^\dagger c_{2\uparrow} + c_{2\uparrow}^\dagger c_{2\uparrow} \bar{n}_{2\uparrow} - \bar{n}_{2\uparrow} \bar{n}_{2\uparrow} \right)$$

Now we need to calculate $\bar{n}_{1\uparrow}$ $\bar{n}_{1\downarrow}$ $\bar{n}_{2\uparrow}$ $\bar{n}_{2\downarrow}$ in a 2 particle system. (we are already neglecting $\uparrow\downarrow$ coherences)

Since we also want to treat only the case $S_z = 0$, we assume $\bar{n}_{i\uparrow} = \bar{n}_{i\downarrow}$. Since the Hamiltonian is invariant under the transformation $1 \leftrightarrow 2$ we also take $\bar{n}_{1\uparrow} = \bar{n}_{2\uparrow} = \bar{n}_{1\downarrow} = \bar{n}_{2\downarrow} = n$. Since

$$\sum_{i\uparrow} \bar{n}_{i\uparrow} = 2 = 4n \quad n = \frac{1}{2}$$

$$H_{\text{HF}} = \epsilon_0 \sum_{i\uparrow} c_{i\uparrow}^\dagger c_{i\uparrow} + t \sum_r (c_{2r}^\dagger c_{2r} + c_{2r}^\dagger c_{1r}) + \\ + \frac{U}{2} \sum_{i\uparrow} c_{i\uparrow}^\dagger c_{i\uparrow} - \frac{U}{2}.$$

This Hamiltonian is easily diagonalized and the eigenvalues are

$$\epsilon_0 + \frac{U}{2} + t \quad \text{for the } r=1 \text{ end } \dagger$$

$$\epsilon_0 + \frac{U}{2} - t \quad \text{for the } r=2 \text{ end } \dagger$$

\Rightarrow the ground state energy in HF:

$$E_g = 2\epsilon_0 + U - \frac{U}{2} + 2t = 2\epsilon_0 + \frac{U}{2} + 2t$$

Let us compare this result with the exact. ($t < 0$)

$$2\epsilon_0 + \frac{U}{2} + 2t \leq 2\epsilon_0 + \frac{U}{2} - \frac{U}{2} \sqrt{1 + \frac{16t^2}{U^2}}$$

$$2|t| \geq \frac{U}{2} \sqrt{1 + \frac{16t^2}{U^2}}$$

$$4|t|^2 \geq \frac{U^2}{4} \left(1 + \frac{16t^2}{U^2}\right) \quad \Rightarrow 0 \geq \frac{U^2}{4}$$

Hartree-Fock over-estimate the ground state energy and is better the smaller the interaction.

11.2 Homogeneous electron gas in the Hartree-Fock approx.

a)

$$\frac{1}{V} \sum_{\mathbf{k}' < k_F} \frac{1}{|\mathbf{k} - \mathbf{k}'|^2} = \frac{1}{V} \int_{\mathbf{k}' < k_F} \frac{d\mathbf{k}'}{(2\pi)^3} \cdot \cancel{L_x L_y L_z} \frac{1}{|\mathbf{k} - \mathbf{k}'|^2} =$$

$$= \frac{1}{(2\pi)^3} \int_0^{k_F} dk' \int_0^{2\pi} d\varphi \int_0^\pi d\theta k'^2 \sin\theta \frac{1}{k^2 + k'^2 - 2kk' \cos\theta} =$$

$$= \frac{1}{4\pi^2} \int_0^{k_F} dk' \int_{-1}^1 d\alpha \frac{k'^2}{k^2 + k'^2 - 2kk'\alpha} = \frac{-1}{4\pi^2} \int_0^{k_F} dk' \frac{k'^2}{2k} \left(\ln |k^2 + k'^2 - 2kk'\alpha| \right) \Big|_{-1}^1$$

$$= \frac{1}{4\pi^2} \int_0^{k_F} dk' \frac{k'}{2k} \ln \left| \frac{(k'+k)^2}{(k'-k)^2} \right| = \frac{1}{4\pi^2 k} \int_0^{k_F} dk' \left[k' \ln |k'+k| - k' \ln |k'-k| \right]$$

$$= \frac{1}{4\pi^2 k} \left\{ \left(\frac{k'^2}{2} \ln |k'+k| \right) \Big|_0^{k_F} - \int_0^{k_F} \frac{k'^2}{2} \cdot \frac{1}{k'+k} - \left(\frac{k'^2}{2} \ln |k'-k| \right) \Big|_0^{k_F} + \int_0^{k_F} \frac{k'^2}{2} \cdot \frac{1}{k'-k} \right\}$$

$$= \frac{1}{4\pi^2 k} \left\{ \frac{k_F^2 - k^2}{2} \ln |k'+k| - \frac{k'^2}{4} + \frac{kk'}{2} - \frac{k'^2 - k^2}{2} \ln |k'-k| + \frac{k'^2}{4} + \frac{kk'}{2} \right\} \Big|_0^{k_F}$$

$$= \frac{1}{4\pi^2 k} \left(\frac{k_F^2 - k^2}{2} \ln \left| \frac{k_F+k}{k_F-k} \right| + kk_F \right) = \frac{1}{2\pi^2} k_F \left(\frac{1}{2} + \frac{1 - (k/k_F)^2}{4(k/k_F)} \ln \left| \frac{1 + k/k_F}{1 - k/k_F} \right| \right)$$

$$b) \sum_{\mathbf{k} < k_F} \frac{k_F^2 - k^2}{kk_F} \ln \left| \frac{k+k_F}{k-k_F} \right| = \frac{V}{(2\pi)^3} \int_0^{k_F} dk \int_0^{2\pi} d\varphi \int_0^\pi d\theta k^2 \sin\theta \frac{k_F^2 - k^2}{k k_F} \ln \left| \frac{k+k_F}{k-k_F} \right|$$

$$= \frac{V}{2\pi^2} \int_0^{k_F} dk \frac{k(k_F^2 - k^2)}{k_F} \ln \left| \frac{k+k_F}{k-k_F} \right| = \frac{V}{2\pi^2} \int_0^{k_F} dk \frac{1}{2k_F} k^2 (k_F^2 - k^2) \cdot \left(\frac{1}{k+k_F} - \frac{1}{k-k_F} \right)$$

$$= \frac{V}{6\pi^2} k_F^3$$

11.3 Thomas Fermi limit in atomic physics

$$V(\bar{r}) = \int d\bar{r}' \frac{e^2}{|\bar{r}-\bar{r}'|} n(\bar{r}') - \frac{Ze^2}{r}$$

a) $n(\bar{r}) = \frac{[2m(\epsilon_F - V(\bar{r}))]^{3/2}}{3\pi^2}$. The reason is that

one can assume a local Fermi energy $\epsilon_F(r) = \epsilon_F - V(\bar{r})$

Now the local density is derived in a canonical way:

$$N = \sum_{\mathbf{k}\sigma} n_{\mathbf{k}\sigma}(\epsilon_{\mathbf{k}}) = 2V \int_0^{k_F} \frac{d\mathbf{k}}{(2\pi)^3} =$$

$$= \frac{1}{\pi^2} V \frac{1}{3} k_F^3 = \frac{V}{3\pi^2} \left(\frac{\hbar^2 k_F^2}{2m} \cdot \frac{2m}{\hbar^2} \right)^{3/2} = \frac{V}{3\pi^2} \left(\frac{2m}{\hbar^2} \epsilon_F \right)^{3/2} \quad \hbar=1$$

$$n = \frac{(2m\epsilon_F)^{3/2}}{3\pi^2}$$

- b) The Fermi energy ϵ_F is uniquely defined all over the space. Far from the nucleus $n(r) \rightarrow 0$ since this is the electronic density and, at the same time $V(r) \rightarrow 0$ since the charge of the nucleus compensates the electronic charge. From a) it follows:

$$\boxed{\epsilon_F = 0}$$

c) It is enough to take the Laplacian of (*)

$$\begin{aligned}\nabla_r^2 V(r) &= \int dr' \nabla_r^2 \frac{e^2}{|r-r'|} n(r') - \nabla_r^2 \frac{ze^2}{r} \\ &= -\int dr' 4\pi \delta(r-r') e^2 n(r') - 4\pi ze^2 \delta(r) \\ &= -4\pi [e^2 n(r) - ze^2 \delta(r)]\end{aligned}$$

$$V(r) = e\phi(r) \Rightarrow -\nabla^2 \phi(r) = 4\pi [en(r) - ze\delta(r)]$$

Thus for $r \neq 0$

$$-\nabla^2 \phi(r) = 4\pi en(r)$$

$$-\nabla^2 V(r) = \left[\frac{1}{r^2} \frac{\partial}{\partial r} r^2 \frac{\partial}{\partial r} \right] e\phi(r) = \frac{4\pi e^2 [2m(\epsilon_F - V(r))]^{3/2}}{3\pi^2}$$

$$-\frac{3\pi}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial V(r)}{\partial r} \right) = 4e^2 [2m(\epsilon_F - V(r))]^{3/2}$$

$$d) \quad V(r) = -\frac{ze^2}{r} \Phi(x) \quad r = z^{-1/3} bx \quad b = \frac{1}{2} \left(\frac{3\pi}{4} \right)^{2/3} a_0 \quad a_0 = \frac{1}{me^2}$$

$$\boxed{\epsilon_F = 0}$$

$$+ \frac{3\pi}{x^2} \frac{z^{2/3}}{b^2} \frac{\partial}{\partial x} \left(x^2 \frac{\partial}{\partial x} \frac{ze^2}{xb} z^{1/3} \Phi(x) \right) = +4e^2 (2m)^{3/2} \frac{z^{3/2} e^3}{b^{3/2} x^{3/2}} z^{1/2} \Phi^{3/2}(x)$$

$$\frac{3\pi}{b^2} \frac{z^{2/3}}{x^2} \frac{\partial}{\partial x} \left(x^2 \frac{\partial}{\partial x} \frac{1}{x} \Phi(x) \right) = \frac{4e^3 (2m)^{3/2}}{b^{3/2}} \frac{1}{x^{3/2}} \Phi^{3/2}(x)$$

$$b^{3/2} = \frac{1}{2^{3/2}} \frac{3\pi}{4} e_0^{3/2} = \frac{3\pi}{4} \cdot \frac{1}{(2me^2)^{3/2}}$$

$$\frac{1}{x^2} \frac{\partial}{\partial x} \left[x^2 \frac{\partial}{\partial x} \left(\frac{1}{x} \phi(x) \right) \right] = \frac{1}{x^{3/2}} \phi^{3/2}(x)$$

$$\frac{1}{x^2} \frac{\partial}{\partial x} \left[x^2 \left(-\frac{1}{x^2} \phi(x) + \frac{1}{x} \frac{\partial}{\partial x} \phi(x) \right) \right] =$$

$$= \frac{1}{x^2} \frac{\partial}{\partial x} \left[-\phi(x) + x \frac{\partial}{\partial x} \phi(x) \right] = x \frac{\partial^2}{\partial x^2} \phi(x) = \frac{1}{x^{3/2}} \phi^{3/2}(x)$$

$$\boxed{\frac{d^2}{dx^2} \phi(x) = x^{-1/2} \phi^{3/2}(x)}$$

$$\lim_{x \rightarrow 0} \Phi(x) = \lim_{r \rightarrow 0} -\frac{r}{2e^2} V(r) = 1 \quad \text{since } u(r) \text{ is not contributing}$$

$$\lim_{x \rightarrow \infty} \Phi(x) = \lim_{r \rightarrow \infty} -\frac{r}{2e^2} V(r) = 0 \quad \text{since } V(r) = 0 \text{ outside the atom.}$$

e)